

Simplified SUSY at Lepton Colliders

Mikael Berggren¹

¹DESY, Hamburg

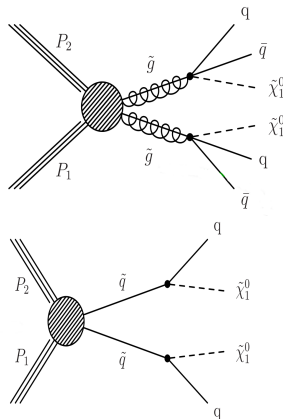
Snowmass Energy Frontier Workshop, BNL , Apr 2013

Outline

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- 2 Simplified SUSY models at LHC
- 3 Simplified SUSY models at ILC
- 4 Simplified SUSY In practice: LEP
- 5 Some results
- 6 Conclusions and Outlook

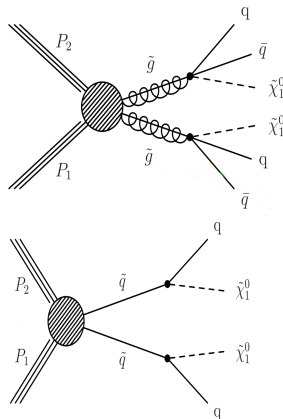
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- **Simplified models** are (very) **special** cases: the produced **SUSY** particle goes **directly** to it's SM partner+MET.
- Either:
 - Production needs a **gluino** and/or **squark** in reach.
- Or: Very special spectra
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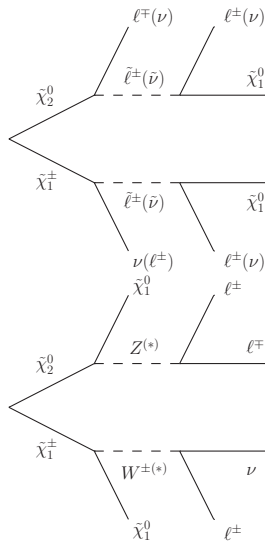


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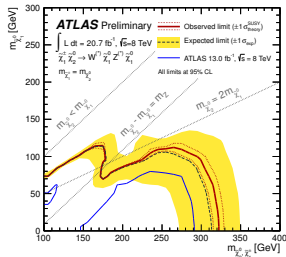
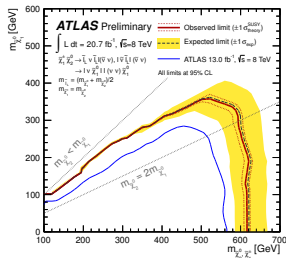
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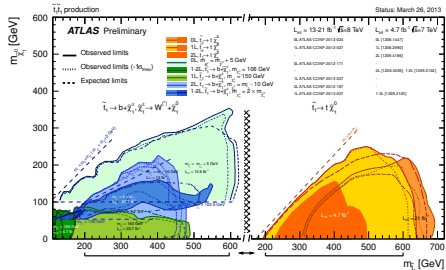
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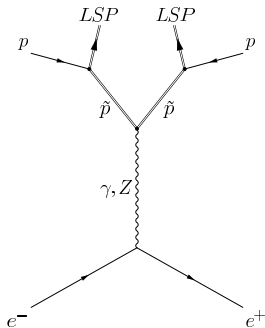
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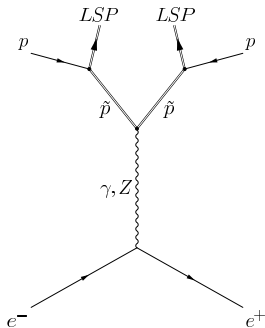
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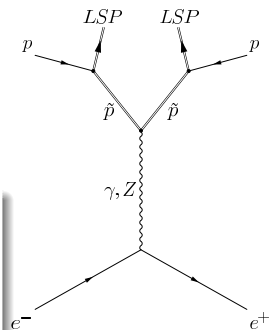
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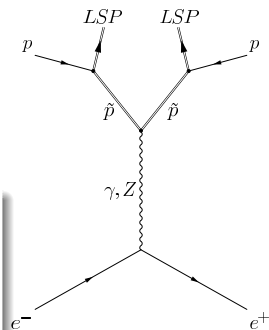
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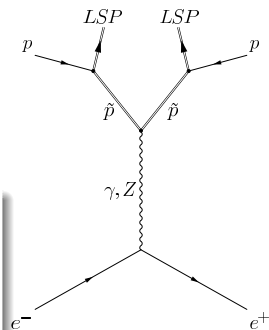
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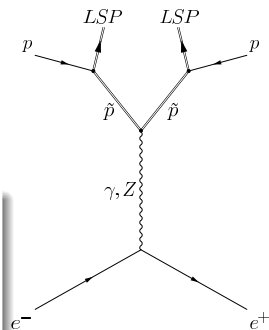
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Simplified SUSY models at ILC: But ?

- RPV:
 - If \exists long-lived charged LSP: Even better
 - If \exists long-lived neutral LSP: same as no RPV.
 - If LSP intermediate: decays in detector, also better.
 - If prompt LSP decay: More complex - combinations of λ , λ' and λ'' constrained by other observations \rightarrow lots of cases, with different signatures. Nevertheless; doable.
- Mixed sparticles:
 - sfermion NLSP: One more parameter. NB: one can't mix away $e^+e^- \rightarrow \tilde{f}\tilde{f}$ completely: Coupling to Z might vanish, but not to γ (exception: \tilde{a})
 - bosino NLSP: Back up one step and evaluate limiting cross-sections instead.
- Very low $\Delta(M)$
 - ISR trick.
 - If $E_{CMS} \gg$ threshold: boost.
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- Side remark: Many open channels, ie if SUSY is main background to SUSY:
 - When data starts coming in, what is first light ?
 - How do we quickly determine a set of approximative model parameters ?
 - What is then the optimal use of beam-time in such a scenario ?
 - And in a staged approach ?
 - Spectrum in continuum vs. threshold-scans?
 - Special points, eg. between $\tilde{\tau}_1 \tilde{\tau}_2$ and $\tilde{\tau}_2 \tilde{\tau}_2$ thresholds.
 - Clean vs. high cross-section.
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- “Parameter” scan:
 - Scan $M_{NLSP} - M_{LSP}$ plane.
 - σ from SUSY-principle and kinematics.
- Do FullSim in $\mathcal{O}(\text{a few})$ points.
- Tune FastSim to these.
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Then:

- If $> 5\sigma$: the point is within **Discovery Reach**.
- If $< 5\sigma$, but $> 2\sigma$: the point is within **Exclusion Reach**.

In practice: LEP

- See <http://lepsusy.web.cern.ch/lepsusy/>
- Sleptons ...
- Squarks ...
- Bosinos ...

These are combined results;
there's more in the results of the
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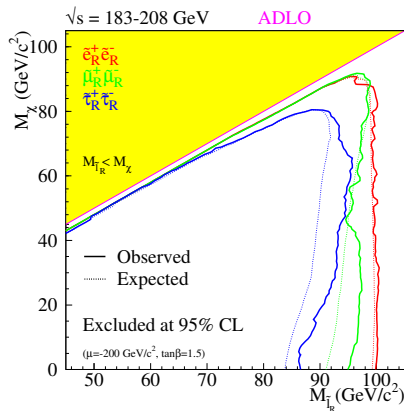
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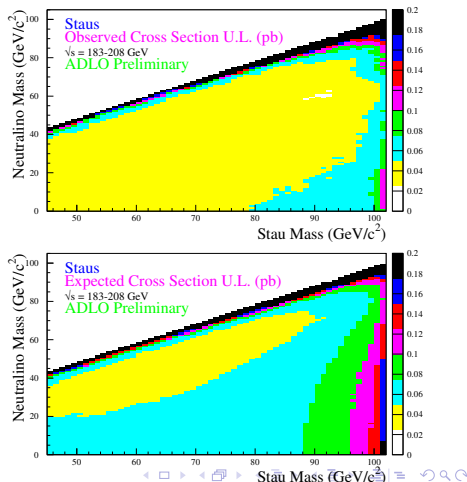


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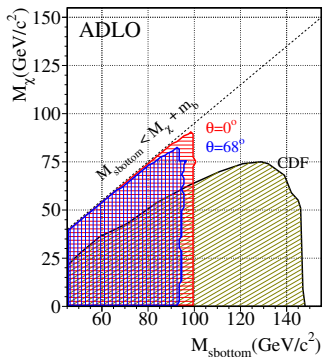


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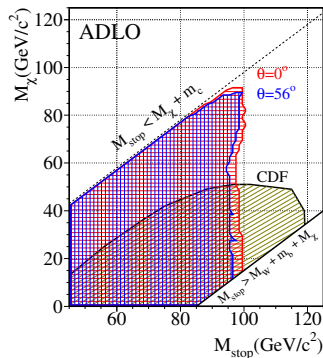


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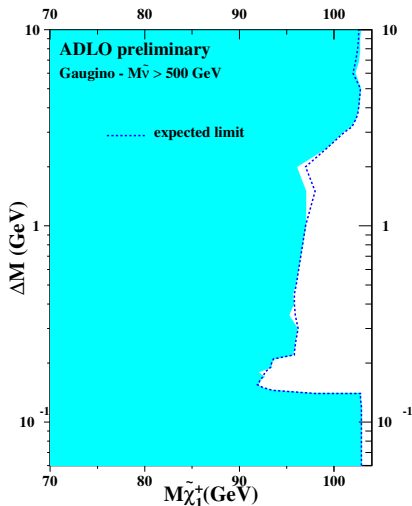


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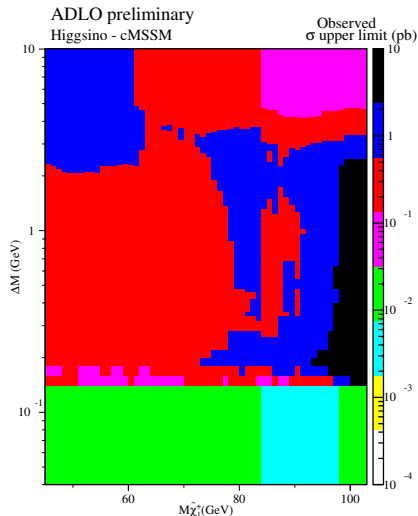


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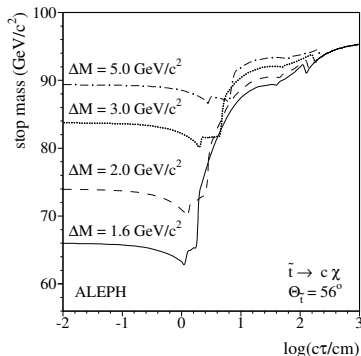
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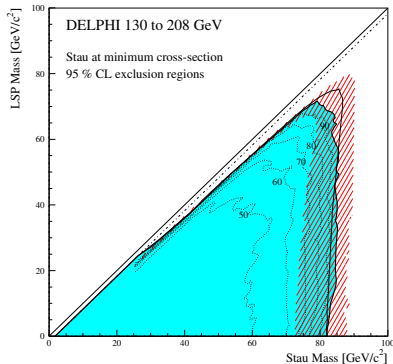


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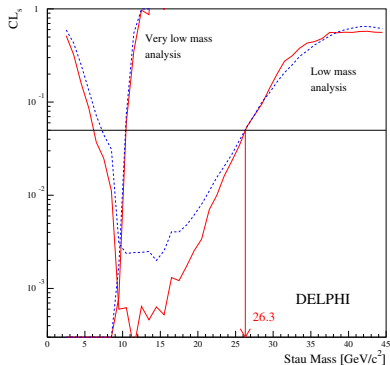


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Bottom line:

VERY hard to wriggle out of this !

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Some results

- Example from existing FullSim and/or FastSim studies:
 - Sleptons, some bosinos in a specific point.
 - Here: $\tilde{e}_R, \tilde{\mu}_R, \tilde{\tau}_1$.
 - Example of the last issue - many open channels.
- Use this to
 - Choose running scenario.
 - Tune FullSim and FastSim to agree ?
- ... then start scanning

Early discovery channel:
cross-section in the pb-range.

- Few simple cuts.
 - $E_{vis} < 400 \text{ GeV}$
 ($= E_{CMS} - 2M_{\tilde{\chi}_1^0 min, LEP}$).
 - 2 charged particles
 - $< 40\%$ of $E_{vis} < \text{below } 30$
degrees.
- Simple observable: E_{vis} : Peak
and width gives $M_{\tilde{e}_R}$ and $M_{\tilde{\chi}_1^0}$.
- See the signal appearing after
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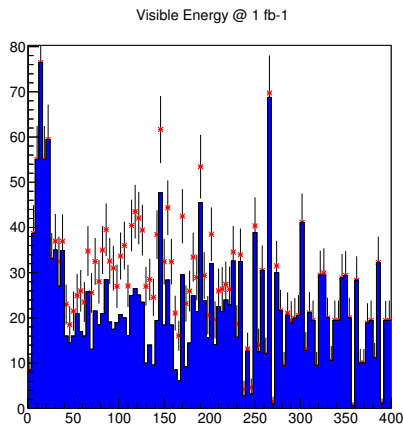
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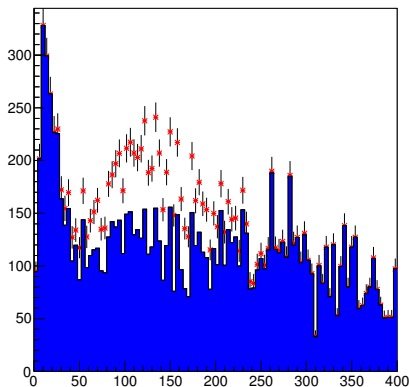
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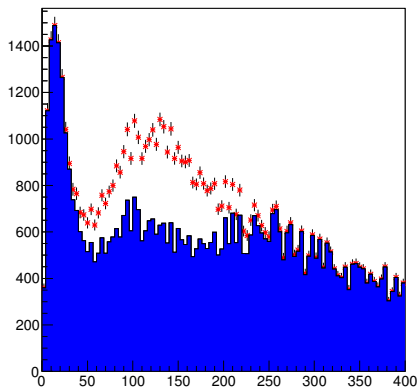
Visible Energy @ 5 fb-1



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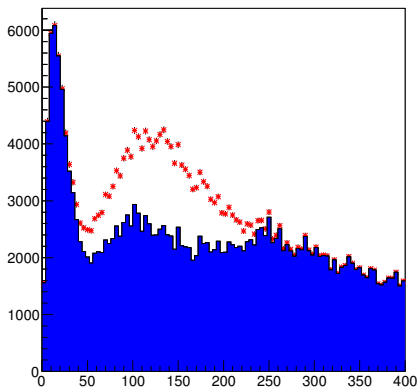
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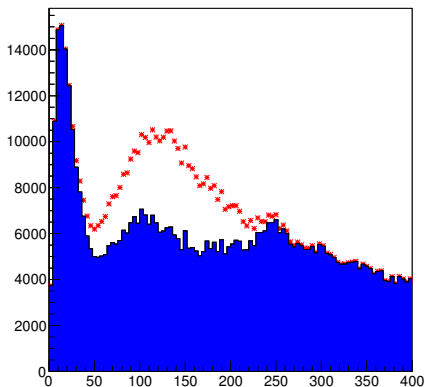
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Visible Energy @ 100 fb-1



Early discovery channel:
cross-section in the pb-range.

- Few simple cuts.
 - $E_{vis} < 400$ GeV
($= E_{CMS} - 2M_{\tilde{\chi}_1^0 min, LEP}$).
 - 2 charged particles
 - $< 40\%$ of $E_{vis} <$ below 30 degrees.
- Simple observable: E_{vis} : Peak and width gives $M_{\tilde{e}_R}$ and $M_{\tilde{\chi}_1^0}$.
- See the signal appearing after
 - 1 fb^{-1}
 - 5 fb^{-1}
 - 25 fb^{-1}
 - 100 fb^{-1}
 - 250 fb^{-1}

Visible Energy @ 250 fb⁻¹

\tilde{e}_R spectrum

- So, within **months** after start-up, we can estimate $M_{\tilde{e}_R}$ and $M_{\tilde{\chi}_1^0}$ to within a **few GeV**.
- Use this knowledge for **better selection** cuts.
- Probably, we have also seen the $\tilde{\mu}_R$.
- ... and that it has \approx the **same mass**. as the \tilde{e}_R

Nets step:

Refine cuts for \tilde{e}_R and $\tilde{\mu}_R$

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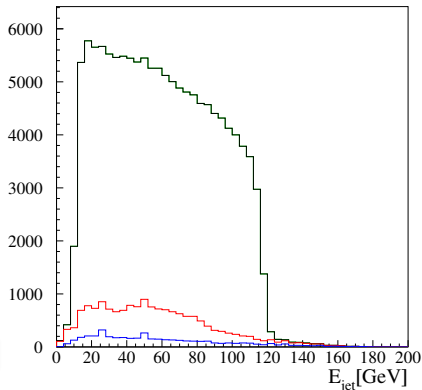
Refine cuts for \tilde{e}_R and $\tilde{\mu}_R$

\tilde{e}_R spectrum

Refine cuts:

- $E_{vis} < 300$ GeV.
- $M_{miss} > 250$ GeV.
- E below 30 degrees < 10 GeV.
- $\cos \theta_{miss} < 0.95$.
- Exactly two opposite charged identified e:s.
- $(E_{jet1} + E_{jet2}) \sin \theta_{acop} > 21$, < 135 GeV.

Efficiency 52 %

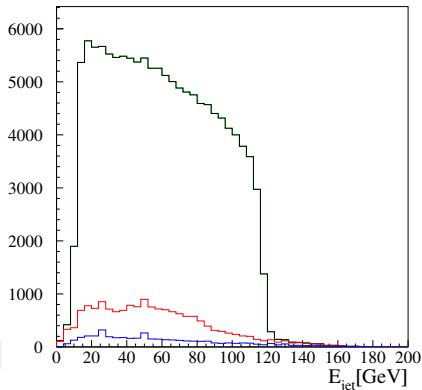


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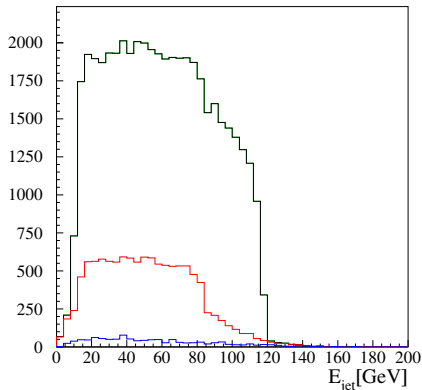
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$\tilde{\mu}_R$ spectrum

Same cuts, but ask for two μ :s instead, ie.:

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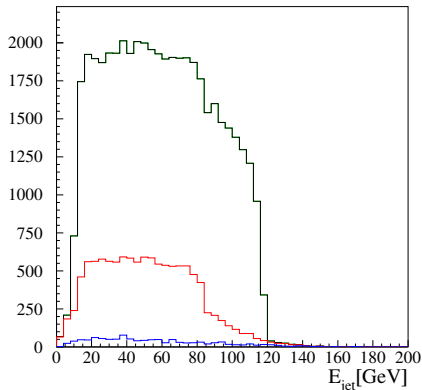


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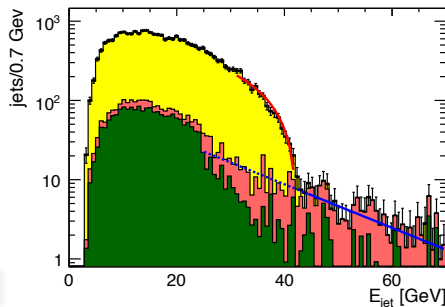


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$\tilde{\tau}_1$ spectrum

- $E_{vis} < 300$ GeV.
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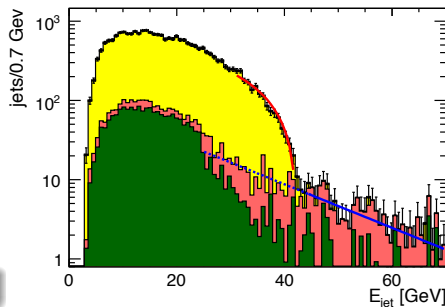
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Conclusions and Outlook

- Simplified methods at hadron and lepton machines are different beasts.
- At lepton machines they are quite model independent, as all possible NLSP's can be exploited in a series 2-dim scans
- This is exactly what was done at LEP, so the procedure is known.
- Now being set up for ILC studies.
- Expect to have a rather complete study before the end of the Snowmass process.

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Work in progress.
Stay tuned.

Thank You !

Backup

BACKUP SLIDES

A New bench-mark point

Remember, apart from naturalness:

- Anomaly in $g - 2$ of the μ : Would prefer a not-too-heavy smuon.
- Dark matter : A WIMP of ~ 100 GeV would be required.
- EW symmetry breaking, coupling constant unification: points to NP at or below 1 TeV
- Suppress the SUSY flavour problem (FCNC:s etc): Heavy 1:st & 2:nd generation squarks would be nice ...
- Other low-energy constrains : $b \rightarrow s\gamma$, $b \rightarrow \mu\mu$, ρ -parameter, $\Gamma(Z)$
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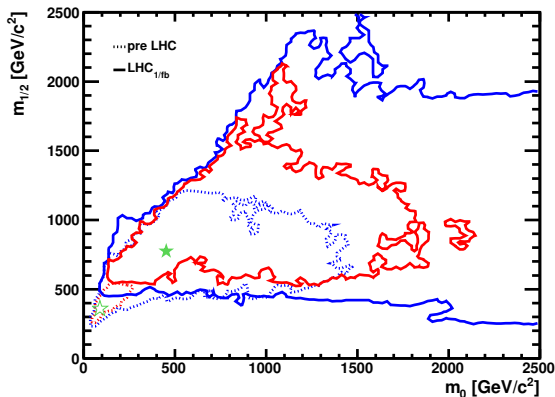
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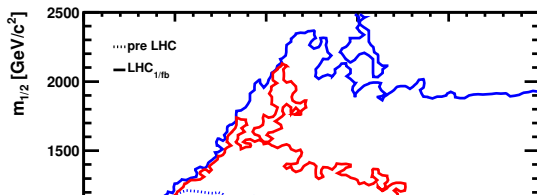
Remember: Without LHC Sps1a' is the best fit!



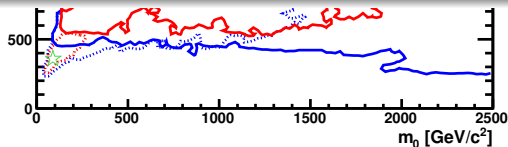
(From Mastercode).

A New bench-mark point

Remember: Without LHC Sps1a' is the best fit!



Can we still get all this with SUSY, without contradicting LHC limits ?!



(From Mastercode).

New points

Can all this be provided by SUSY ? Yes, sure !

Take old ILC favourite benchmark SPS1a, and make the TDR4 point
(see Baer&List arXiv:1205.6929v1

New points

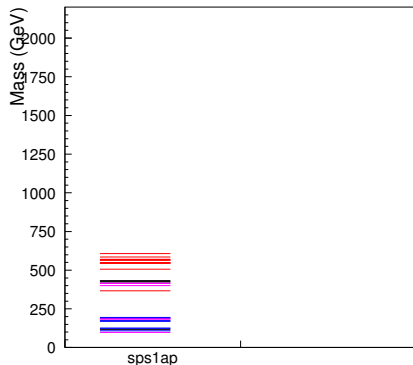
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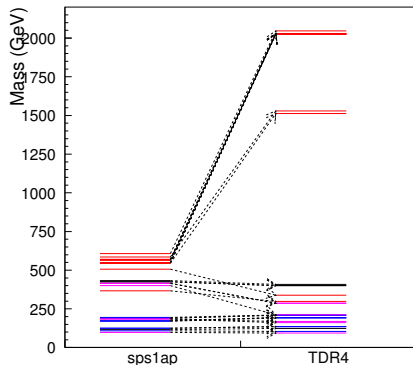
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SPS1a: mSUGRA

- 5 parameters.
- One gaugino parameter
- One scalar parameter

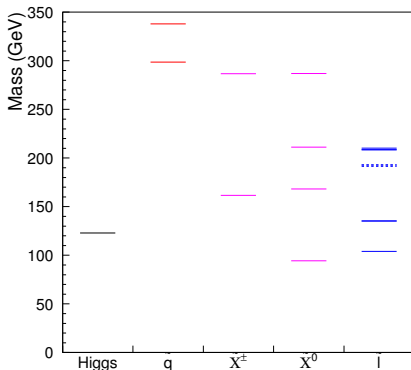
TDR4: Phenomenological SUSY

- 11 parameters.
- Separate gluino
- Higgs, un-coloured, and coloured scalar parameters separate

Parameters chosen to deliver all constraints, \approx **same ILC accessible spectrum** \Rightarrow old analyses **still valid !**

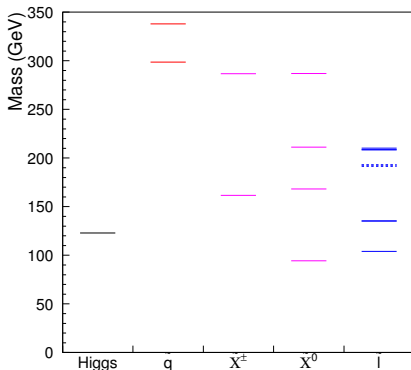
Features of TDR 4

- The $\tilde{\tau}_1$ is the NLSP.
- For $\tilde{\tau}_1$: Small Δ_M , $\gamma\gamma$ - background
- For $\tilde{\tau}_2$: $WW \rightarrow l\nu l\nu$ - background \Leftrightarrow Polarisation.
- $\tilde{\tau}$ NLSP $\rightarrow \tau$:s in most SUSY decays \rightarrow SUSY is background to SUSY.
- For $\text{pol}=(-1,1)$: $\sigma(\tilde{\chi}_2^0 \tilde{\chi}_2^0)$ and $\sigma(\tilde{\chi}_1^+ \tilde{\chi}_1^-) = \text{several hundred fb and}$ $\text{BR}(X \rightarrow \tilde{\tau}) > 50\%$. For $\text{pol}=(1,-1)$: $\sigma(\tilde{\chi}_2^0 \tilde{\chi}_2^0)$ and $\sigma(\tilde{\chi}_1^+ \tilde{\chi}_1^-) \approx 0$.

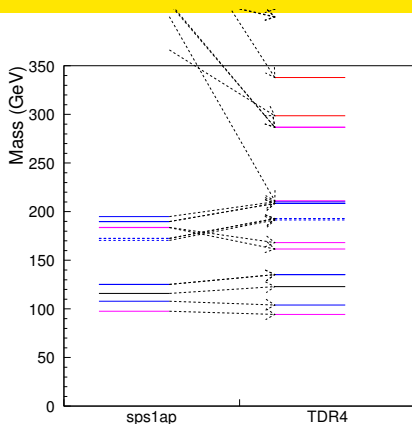


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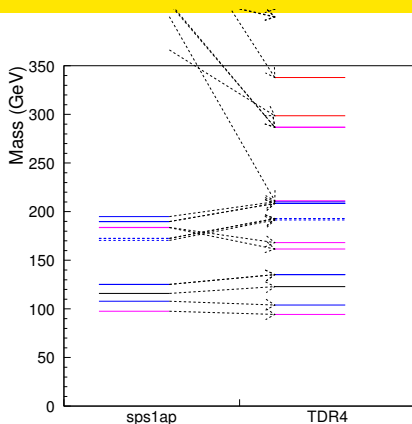


Differences TDR4 - SPS1a'



- All bosinos
- M_h OK
- $\tilde{\ell}_L \rightarrow \tilde{\chi}_0^0 \ell$ at 30-40 % BR.
- $\tilde{\chi}_4^0$ and $\tilde{\chi}_2^\pm$ too heavy
- M_h too small
- $\tilde{\ell}_L \rightarrow \tilde{\chi}_0^0 \ell$ at ~ 95 % BR.

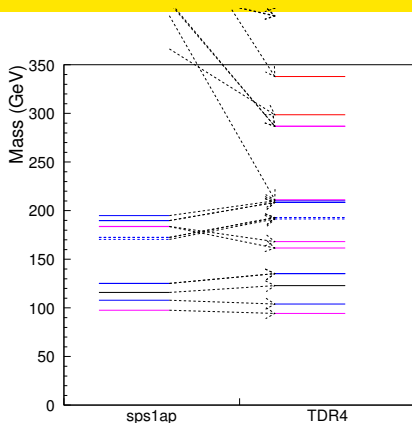
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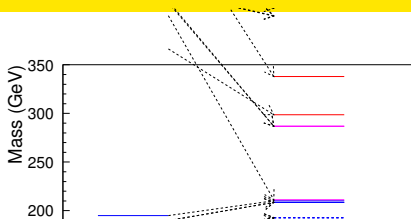
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Bottom line

Even more open channels

More complicated topologies

We plan to check how close TDR4 is to the “best fit” (with **fittino**)



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Analysis

Disclaimer

- Very preliminary
- Mostly taken over SPS1a' analyses: Guaranteed to have bad efficiency for heavier states, due to the increase of cascade decays (mostly ignored in Sps1a')

Take over SPS1a' (Phys.Rev.D82:055016,2010, Nicola's thesis,...)

Lighter sleptons

Use the polarisation (0.8,-0.3) of the data to reduce bosino background. Assumed to be 50 % of all data.

From decay kinematics:

- $m_{\tilde{\ell}}$ and $M_{\tilde{\chi}_1^0}$ and end-points of spectrum = $E_{\ell, \min(\max)}$.
- For $\tilde{\tau}_1$: other end-point hidden in $\gamma\gamma$ background: Must get $M_{\tilde{\chi}_1^0}$ from other sources. ($\tilde{\mu}$, \tilde{e} , ...)

$m_{\tilde{\ell}}$ also from cross-section:

- $\sigma_{\tilde{\ell}} = A(\theta_{\tilde{\ell}}, \mathcal{P}_{beam}) \times \beta^3/s$, so
- $m_{\tilde{\ell}} = E_{beam} \sqrt{1 - (\sigma s/A)^{2/3}}$: no $M_{\tilde{\chi}_1^0}$!

From decay spectra:

- \mathcal{P}_{τ} from exclusive decay-mode(s): handle on mixing angles $\theta_{\tilde{\tau}}$ and $\theta_{\tilde{\chi}_1^0}$

Topology selection

Take over SPS1a' $\tilde{\tau}$ analysis principle

$\tilde{\ell}$ properties:

- Only two particles (possibly τ :s:s) in the final state.
- Large missing energy and momentum.
- High Acolinearity, with little correlation to the energy of the τ decay-products.
- Central production.
- No forward-backward asymmetry.

+ anti $\gamma\gamma$ cuts (see backup)

Select this by:

- Exactly two jets.
- $N_{ch} < 10$
- Vanishing total charge.
- Charge of each jet = ± 1 ,
- $M_{jet} < 2.5 \text{ GeV}/c^2$,
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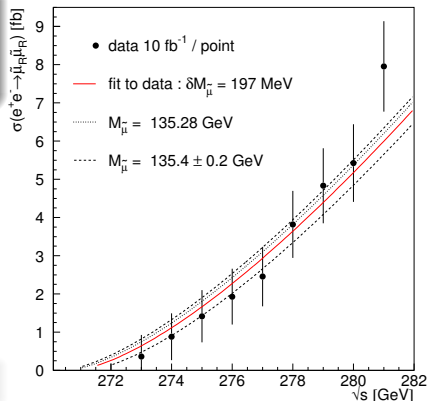
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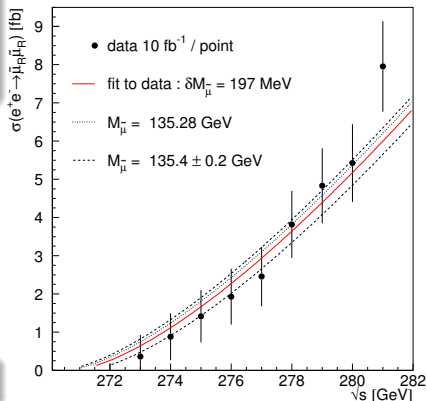
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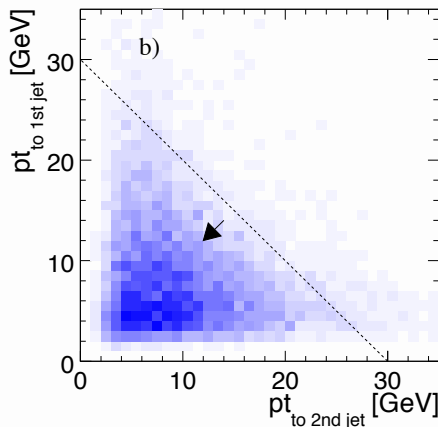
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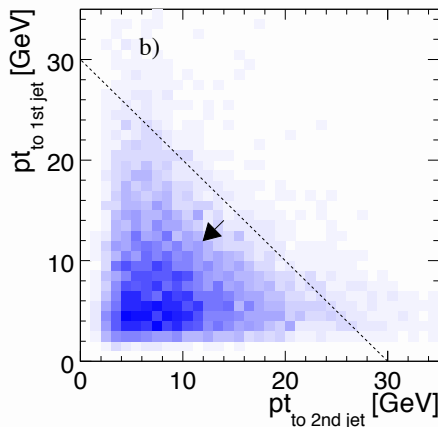
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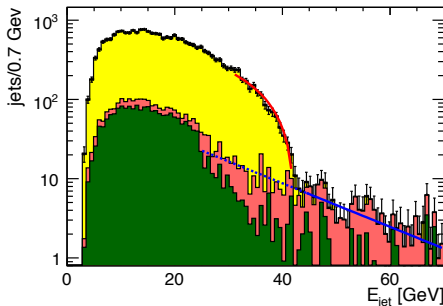


Fitting the $\tilde{\tau}_1$ mass (SPS1a')

- Only the upper end-point is relevant.
- Background subtraction:
 - Important SUSY background, but region above 45 GeV is signal free. Fit exponential and extrapolate.
- Fit line to (data-background fit).

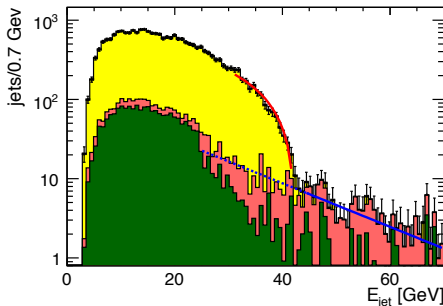
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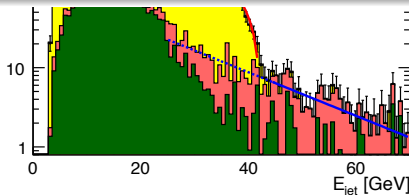
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Results from cross-section for $\tilde{\tau}_1$

$$\Delta(N_{\text{signal}})/N_{\text{signal}} = 3.1\% \rightarrow \Delta(M_{\tilde{\tau}_1}) = 3.2 \text{ GeV}/c^2$$



First look at Heavier sleptons ($\tilde{\mu}_L$)

Remember

demanding exactly 2 objects kills 90 % of the signal in TDR4, due to cascaded decays !

- Same cuts as for $\tilde{\mu}_R$, and
- anti-WW likelihood, take over from SPS1a'
- select using other particle:
 $p(\text{other } \mu) > 120 \text{ GeV}$.

Efficiency 1.5 % (!), $S/B = 0.2$.

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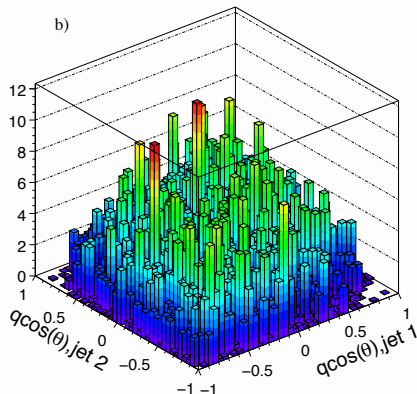
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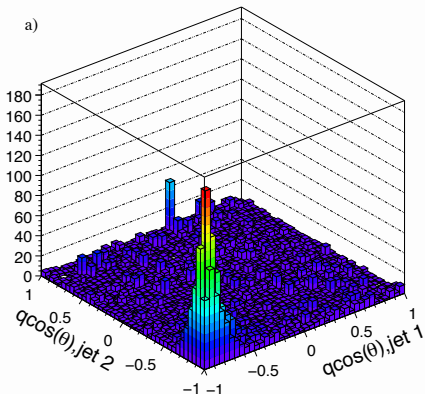
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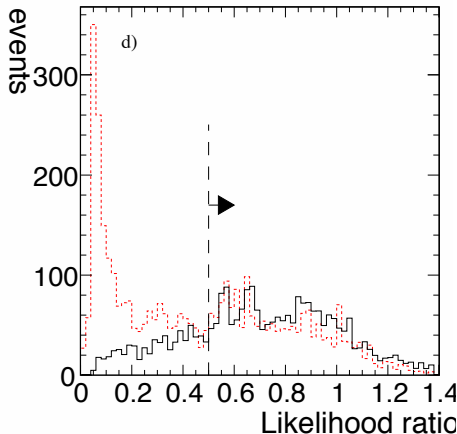
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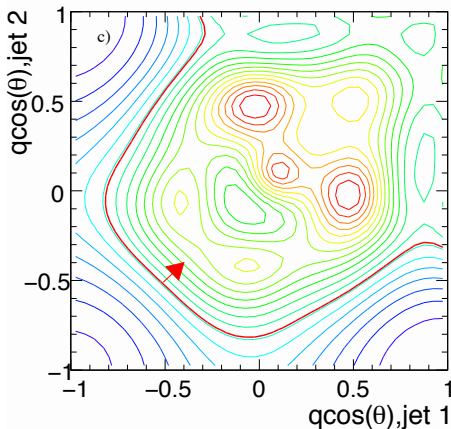
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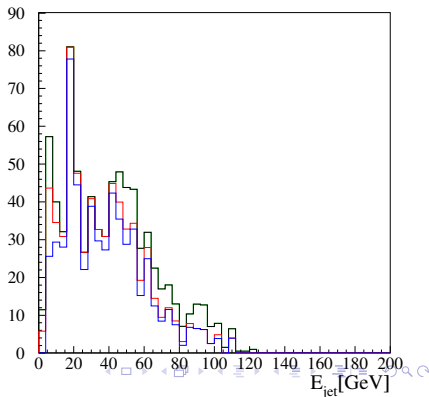
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